



# Xiamen Valence Bond

## An *ab initio* Non-orthogonal Valence Bond Program

Version 3.1

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Last updated on May 5, 2023

# Referencing for XMVB

You need to cite the references of XMVB as following formats when the calculations are proceeded with XMVB program.

A) J. Chem. Phys. format:

The *ab initio* Valence Bond calculations are performed with the XMVB program.<sup>1</sup>

1. (a) Z. Chen, F. Ying, X. Chen, J. Song, P. Su, L. Song, Y. Mo, Q. Zhang and W. Wu, *Int. J. Quantum. Chem.*, 2015, 115, 737 (b) L. Song, Y. Mo, Q. Zhang, W. Wu, *J. Comput. Chem.* 2005, 26, 514.

B) American Chemical Society format:

The *ab initio* Valence Bond calculations are performed with the XMVB program.<sup>1</sup>

1. (a) Chen, Z.; Ying, F.; Chen, X.; Song, J.; Su, P.; Song, L.; Mo, Y.; Zhang, Q.; Wu, W., *Int. J. Quantum. Chem.*, **2015**, 115, 737. (b) Song, L.; Mo, Y.; Zhang, Q.; Wu, W. *J. Comput. Chem.* **2005**, 26, 514.

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# Availability of XMVB package

**XMVB** version 3.0 is a program that performs *ab initio* valence bond calculations. **XMVB** version 3.0 will be distributed from Xiamen University for free. The principal user will be asked to fill a license agreement and send the signed agreement to:

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# Chapter 1

## INTRODUCTION

### 1.1 What is XMVB?

Xiamen Valence Bond (XMVB) is a quantum chemistry program for performing electronic structure calculations based on the non-orthogonal Valence Bond methods.

In the current arena of theoretical and computational chemistry, molecular orbital (MO) based methods undoubtedly are enjoying the overwhelming dominance, largely due to their high computational efficiencies. Apart from the MO theory, however, valence bond (VB) theory remains the widespread conceptual matrix for chemists, thus the development of novel algorithms and practical programs for the *ab initio* VB theory have been an endeavor for many of us.

The VB project at Xiamen University was initiated in 1986, when a spin-free form of VB method was independently proposed. In 1989, we wrote a simple VB code and applied it to H<sub>3</sub> molecule. The systematic development of a complete and efficient VB code nevertheless gained the momentum in 1992, when an algorithm based on the left coset decomposition of the symmetric group was proposed and programmed.<sup>[1,2]</sup> In 1995, we further developed an algorithm called the paired-permanent-determinant (PPD) approach,<sup>[3-6]</sup> which is more efficient for systems of many covalent bonds than the traditional Slater expansion algorithm, on which our previous VB code, Xiamen-99, was based. In the last twenty years, we continued optimizing the code and implementing some newly developed VB approaches into the code. To distinguish our VB code from our city whose name is Xiamen, the code was renamed as XMVB<sup>[7]</sup> in 2002, beginning with Version 1.0. After ten years, Version 2.0 was released. XMVB 2.1 was released in 2015. XMVB 3.0 was released in 2017. Now, a new version of XMVB, XMVB 3.1, is released.

The motivation of the development of the XMVB package is to provide chemists with a well-designed and efficient program to perform both classical and modern valence bond calculations at *ab initio* level. To gain chemical lucidity from VB theory, the XMVB package uses Heitler-London-Slater-Pauling (HLSP) functions as state functions. The superposition of a set of HLSP functions result in the wave function for a system. VB orbitals may be strictly localized, delocalized, or bonded-distorted (semi-localized), depending on particular applications.

Currently, there are 2 mechanisms of distribution available for users: module and stand-alone package. The module distribution is a module for the GAMESS-US package implemented with Fortran 77; while the stand-alone distribution obtains integrals from utility PREINT

and performs VB computations as a stand-alone program. The syntaxes of XMVB input files are the same for both distributions. Incorporating our VB code into the GAMESS-US package allows more hybrid VB methods, such as DFVB, VBPCM, and VBEFP, to be performed. The stand-alone package is implemented with Fortran 90 and parallelization has been introduced. The users may choose which version to use by themselves.

## 1.2 What's New in Version 3.1?

Compared with XMVB 2.0, 2.1 and 3.0, the following features are available in XMVB 3.1:

- MPI is implemented in XMVB. MPI+OpenMP for multi-node parallelization is now available for following functions:
  1. VBSCF with ISCF=1/2/5 and tensor-based VBSCF
  2. BOVB with ISCF=1/2
  3. VBCIS and VBCISD
- Symmetry is supported for VB computation.

## 1.3 What Can Be Done with XMVB?

- Following ab initio VB methods are implemented in XMVB:
  - Valence Bond Self-Consistent Field (VBSCF);
  - Breathing Orbital Valence Bond (BOVB);
  - Valence Bond Configuration Interaction (VBCI);
  - Valence Bond Perturbation Theory (VBPT2);
  - Density Functional Valence Bond (DFVB);
  - Valence Bond Polarizable Continuum Model (VBPCM);
  - Valence Bond Effective Fragment Potential (VBEFP);
  - Combined Valence Bond Effective Fragment Potential Polarizable Continuum Model (VBEFP/PCM).
- Many-electron wave function of computed system is written in terms of HLSP functions or Slater determinants. Both of structural coefficients and weights are provided.
- The form of VB orbitals is flexible; they may be HAOs, BDOs, or OEOs, which are strictly localized, semi-localized, or fully delocalized.
- Following population analysis properties of wave function are computed:
  - Mulliken and Löwdin atomic charges;
  - Atomic spin density polarization;
  - Mayer's bond order;
  - Dipole moments.
- Interface files for other programs, i.e. AIM2000 and MOLDEN, are provided for computing more properties or viewing VB orbitals.



# Chapter 2

## INSTALLATION

Both distributions of XMVB are currently available for LINUX platform. 1.5GB RAM is required. Followings are the instructions for installation. Note that the source code will NEVER be released to the users. Only compiled object file or executable files are available for users.

### 2.1 For Module Distribution

To build the module distribution, the user may need a library file “libgameess-xmvp.a” which can be found in the package they obtained. The user may also need to prepare LIBCINT library file “libcint.a” which is needed for some functions in XMVB. After that, copy these two library files to “object” directory in GAMESS-US. The user also need to make the following modifications to linke XMVB in to GAMESS-US package. We assume that the user will run GAMESS-XMVB with AMD-64 Linux, which corresponds to “linux64” architecture of GAMESS-US, and the target is “sockets”. All operation is proceeded in the root directory of GAMESS-US.

1. Open “ddi/compddi” and modify

```
set DDI_COMM = '-DDDI_SOC'
```

to

```
set DDI_COMM = '-DDDI_SOC -mcmmodel=large'
```

then run

```
cd ddi && ./compddi
```

to recompile DDI.

2. Open “comp”, find the entry for “linux64”, then modify

```
set EXTRAOPT="" "
```

to

```
set EXTRAOPT="-mcmmodel=large"
```

in the branch of “gfortran”, then run

```
./compall
```

to recompile all modules.

3. Open “lkd”, modify

```
set XMVB=false
```

to

```
set XMVB=true
```

to activate linking XMVB, then modify

```
set VBOBJ='xmvb.o'
```

to

```
set VBOBJ='libgams-xmvb.a libcint.a /path/to/lapack /path/to/blas'
```

where **/path/to/lapack** is the location of LAPACK library file and the same to **/path/to/blas**. The location of LAPACK and BLAS may be omitted if MKL is used.

4. Open “lkd”, modify

```
set LDOPTS=' '
```

to

```
set LDOPTS='-fopenmp -mcmode=large'
```

in the branch of “gfortran”, then run

```
./lkd gams xmvb
```

to start linking. After that, an executable file “gams.xmvb.a” including the whole capability of XMVB will be generated.

5. Before running GAMESS-XMVB, don't forget to append line

```
setenv VBINP $USERSCR/$JOB.xmi
```

in the file “gms-files.csh”. The variable **\$USERSCR** can be replaced with other values such as “\$SCR” defined in “rungms”.

## 2.2 For Stand-Alone Distribution

The stand-alone distribution is released as a package of compiled executable files. To install the stand-alone distribution, the users should:

- A Unpack the compressed tar file by using the following command,

```
tar xvfz xmvb.tar.gz
```

Once the file is unpacked successfully, a new directory **xmvb/** will be created.

- B Set the VBDIR environment variable to the location of XMVB package and append **xmvb/** to your \$PATH. The variable VBDIR is essential for PPD algorithm and utility PREINT(see Section 6.2).

# Chapter 3

## RUNNING A JOB

An XMVB job consists of the following two steps:

- Step 1. Prepare the integrals of primitive basis functions and the MO orbitals of the system. An input file including molecular information and basis set information is needed. Files "1e.tem", "2e.tem" and "orb.mo" will be generated after this step.
- Step 2. Run XMVB calculation. A file with extension name ".xmi"(XMI file, see chapter 4) is needed here.

The programs, utilities and files needed in these 2 steps differ in different distributions. The details are shown below.

### 3.1 For Module Distribution

The two steps of an XMVB job in module distribution are done in one shot with command **rungms job [VERNO] [1]**

after all files are prepared.

*Step 1* is done in GAMESS program with a GAMESS input file including line "\$CONTRL VBTYP=XMVB \$END". XMVB calculation will be proceeded automatically following the completion of *step 1*.

Tips:

1. By default, XMVB guess and input files will not be copied into the \$SCR directory, which is the real directory where GAMESS- US is doing the calculation. Please copy these files into your \$SCR directory *BEFORE* the calculation. Otherwise an error may occur.
2. After the calculation, the integral files and XMVB outputs will also be left in \$SCR. Integral files may be used for XMVB calculations with stand-alone XMVB program. Please remember to get them if you need them later for stand-alone XMVB.

### 3.2 For Stand-Alone Distribution

With stand-alone distribution used, step 1 is done with program PREINT(see Section 6.2). After obtaining the integral, create an input file file.xmi for your job (for details, see Chapter 4). Then run the XMVB job by typing command

**xmvp file.xmi**

Then an output file with name file.xmo will be obtained. For use XMVB in parallelization, see section 3.3.

### 3.3 Parallelization in stand-alone XMVB

MPI and OpenMP are two popular protocols for parallelization. MPI will create a bunch of processes and the data is synchronized by message-passing, and is useful for multi-node parallelization in distributed systems, e.g. clusters. OpenMP creates a bunch of lightweight threads under a process and is useful for parallelization in single node. The MPI+OpenMP strategy may take the advantages of both protocols to provide a highly-efficient parallelization with multiple nodes. With  $N_{\text{mpi}}$  MPI processes and  $N_{\text{omp}}$  OpenMP threads for each process, the total number of CPU cores consumed for a parallel job  $N_{\text{cpu}}$  is

$$N_{\text{cpu}} = N_{\text{mpi}} * N_{\text{omp}}$$

Currently such strategy is supported only by stand-alone XMVB. To use MPI in XMVB, the user may have MPICH installed in the computer. The version of MPICH to compile XMVB is 3.3.2.

#### 3.3.1 MPI parallelization

To run XMVB with MPI parallelization, the user need to type the command

```
mpirun -np np xmvp file.xmi
```

to start a computation with  $np$  processes. For more information about running parallel computation with MPICH, please refer to the [MPICH documentation](#).

#### 3.3.2 OpenMP parallelization

The OpenMP parallelization does not need to change the command of running XMVB jobs, as shown in 3.2. The number of CPU cores for the parallelization can be controlled by setting the variable OMP\_NUM\_THREADS

```
export OMP_NUM_THREADS=N
```

if you just need a certain number  $N$  of CPUs for your job. If the variable is absent, OpenMP parallelizes the XMVB computation with *ALL* processors in your computer.

For large systems, OpenMP parallelization may proceed a strange "segmentation fault". This is because the stack size of threads is not large enough. This can be avoided by setting the stack size to a certain number to avoid this error. In OpenMP parallelization, the stack size of master and slave threads are set in different ways. The stack size of master thread is set by command *ulimit* as shown

```
ulimit -s stack_size
```

The default stack size is 8192. Setting a larger value or simply

```
ulimit -s unlimited
```

The stack size of slave threads are controlled by environment variable  $\$OMP\_STACKSIZE$ . Following command will set the stack size of each slave threads to 1GB

```
export OMP_STACKSIZE=1G
```



# Chapter 4

## INPUT FILE

### 4.1 Syntax of XMI File

An ".xmi" input file describes the information of the XMVB computation, containing up to 7 sections as "CTRL", "BFI", "STR", "FRAG", "ORB", "AIM", and "GUS". The following are the format and example of "xmi" file:

```
$CTRL
nmul=val1, nstr=val2, norb=val3, ...
$END
$STR
ntstr(1,1), ntstr(2,1), ..., ntstr(nel,1)
ntstr(1,n), ntstr(2,n), ..., ntstr(nel,n)
$END
$BFI
nfroz nbas
List of frozen MOs
List of basis functions
$END
$FRAG
nf(1), nf(2), ..., nf(N)
lf(1,1), lf(2,1), ..., lf(nf(1),1)
...
lf(1,n), lf(2,n), ..., lf(nf(n),n)
$END
$ORB
max(1), max(2), ..., max(val3)
nvic(1,1), nvic(2,1), ..., nvic(max(1),1)
...
nvic(1,n), nvic(2,n), ..., nvic(max(n),n)
$END
$AIM
WFN file name
$END
```

```

$GUS
VBO(1) MO for VBO(1)
...
VBO(n) MO for VBO(n)
$END

```

A minimal example of XMVB input file is shown below:

```

VBSCF/3-21G Molecule: HF
$ctrl
nstr=3
$end
$str 1 1 2 3 ; 3 structures
1 1 2 2
1 1 3 3
$end
$orb
4 4 2
3-6
4 3 5 6
1 2
$end

```

#### NOTE FOR INPUT:

- There shall not be more than 400 characters in a line.
- The first line is the explanation line.
- Anything after ";" or "#" in a line will be regarded as comment. If they appear in the same line, the first appearing will be the dominance.
- Section starts with \$CTRL, \$BFI, \$STR, \$FRAG, \$ORB, \$AIM, \$GUS, \$SCF, \$GEO and end with \$END. Don't write anything else in begin or end line.
- All sections except \$CTRL can be optional based on a user's choice of keywords.
- Sections can be placed in any order.
- Abbreviation is available in the whole input file. For example, in the above example, string "4 4 2" may be written equivalently as "4\*2 2"; "1 1 2 2" may be written equivalently as "1:2"; "3-6" means "3 4 5 6".

## 4.2 \$CTRL Section

The \$CTRL section contains the information of how a job is performed. The input format is **name=value** or **name=option**, except for the keywords which need no values or options. <enter> and <space> are used to separate keywords. If a keyword accepts several options in a time, the options are separated with ",".



### 4.2.1 Keywords for Global Control

#### **BPREP:**

This keyword initiates an integral transformation from primitive basis functions to VB basis functions with \$BFI (see below) needed. The transformation may freeze core orbitals, remove some primitive basis functions which are not involved in VB calculation, and hybridize basis functions. XMVB will use primitive basis functions without transformation if this keyword is missing.

#### **NOTE:**

This keyword cannot be used together with ORBTYP=HAO or GUESS=MO (see below).

#### **EPG=*n*:**

Set the convergent criterion of energy to  $10^{-n}$ . Default value is 7.

#### **GPG=*n*:**

Set the convergent criterion of gradient. Floating point number is inputted. Default value is 2.D-3 for ISCF=1/3, 1.D-3 for ISCF=2/5, and 1.D-4 for ISCF=6.

#### **ITMAX=*n*:**

*n* is the maximum number of iterations. Default value is 200.

#### **NMUL=*n*:**

*n* is the spin multiplicity ( $2S + 1$ ) of system. Default value is 1, which means singlet state.

#### **NAO=*m*:**

*m* is the number of active VB orbitals whose occupation number varies in the structures. NAO is required if keywords *STR* or *ISCF*=5 (see below) is specified.

#### **NAE=*n*:**

*n* is the number of active VB electrons which occupy the active orbitals. NAE is required if keywords *STR* or *ISCF*=5 (see below) is specified.

#### **NSTR=*n*:**

*n* is the number of VB structures (or determinants). This keyword can be omitted if *STR* (see below) is assigned.

#### **STR=options:**

This keyword generates VB structures automatically and hence *NSTR* and the \$STR section are not needed. This keyword requires NAO and NAE (see above) to declare the active space. Users may use one or several of the following options:

COV: Covalent structures will be generated.

ION[(*n-m*)]: Ionic structures will be generated. A simple ION will generate all ionic structures; ION(*n,m*) will generate only the *n*-th and *m*-th order ionic structures and ION(*n-m*) will generate ionic structures from the *n*-th to the *m*-th order.

FULL: All VB structures will be generated.

By default, VB structures will not be generated automatically and *NSTR* and *\$STR* are essential and neither *NAO* nor *NAE* is necessary.

**FIXC:**

Request to fix structure coefficients for VB structures. In VB theory, the coefficients are obtained by solving the secular equation  $HC = SCE$ . For some special purposes, one may want to fix the coefficients. The coefficients are inputted following the corresponding VB structures. For example, the following input will constrain the coefficients of the three VB structures to be 1.0:0.5:0.5

```
$STR
1 2 1.0
1 1 0.5
2 2 0.5
$END
```

**GROUP=EXP:**

Divide VB structures into groups according to the expression *EXP*. An expression with  $n$  structures divided into  $m$  groups can be expressed as:

$$\dots, S_{i1}, \dots, \dots, S_{j2}, \dots, \dots, \dots, S_{nm} \dots$$

Here  $S_{i1} \dots S_{nm}$  are the structure numbers, a comma "," is used to separate the structures numbers in the same group, and two commas "," is used to separate different groups. Coefficients of structures should be given in *\$STR*, similar to **FIXC**. The ratio of VB structures within the same group will be fixed, as introduced in **FIXC**. The coefficients of VB structures in different groups will not be fixed and shall be optimized by solving secular equation. Following is an example:

```
$CTRL
NSTR=3 GROUP=1,,2,3
$END
$STR
1 2 1.0 # S1
1 1 0.5 # S2
2 2 0.5 # S3
$END
```

The above example divide 3 VB structures into 2 groups:

Group 1.  $G_1 = S_1$

Group 2.  $G_2 = 0.5(S_2 + S_3)$

Hence a 3 structure problem becomes a 2 "structure" problem:

$$\Psi = C_1 G_1 + C_2 G_2$$

where  $C_1$  and  $C_2$  are coefficients of  $G_1$  and  $G_2$  obtained by solving secular equation. The final wave function can be expressed as

$$\Psi = C_1 S_1 + \frac{C_2}{2} S_2 + \frac{C_2}{2} S_3$$

**NSTATE=n:**

Energy, coefficients and weights of structures for the  $n$ -th excited state, rather than for the

ground state, will be calculated and printed out. The values of  $n$  can be:

- 0: The ground state.(Default)
- $n$ : The  $n$ -th excited state.

NOTE:

- a. VB orbitals are optimized by minimizing the energy of required state. When the  $n$ -th excited state is requested, the  $(n + 1)$ -th root will be chosen as the  $n$ -th excited state when solving the secular equation. Thus,  $n$  must be smaller than the number of structures.
- b. For VBCI calculations, *NSTATE* can be only 0 or 1.

**IPRINT= $n$ :**

Printing levels for XMVB. The available levels are:

- 1: Initial guess, energy, coefficients, weights, and orbitals will be printed. This is the default printing level.
- 2: All contents in IPRINT=1, Hamiltonian and overlap matrices in terms of VB structures, and population analysis will be printed.
- 3: All contents in IPRINT=2, density matrix and orbital overlap matrix will be printed.

**SORT:**

Sort the VB structures in descending order according to coefficients.

**CTOL= $tol$ :**

Set the Coefficient TOLerance when printing coefficients and weights of VB structures. Only the coefficients and weights of VB structures whose absolute values of coefficients are not smaller than tolerance  $tol$  will be printed. The default tolerance is 0, which means all structures will be printed.

NOTE:

The tolerance  $tol$  is a real parameter. For instance,

**ctol=0.01**

It means that only structures whose absolute values of coefficients larger than or equal to 0.01 will be printed. For VBCI this keyword is not functioning.

**CICUT= $n$ :**

Set cut threshold to  $10^{-n}$  for CI configurations. The contribution of a CI configuration is estimated by perturbation theory. If the contribution is less than the threshold, the configuration will be discarded. This will reduce the computational effort for CI calculations. Recommended values are 5 or 6. Default value is 0 (no cut).

**NCOR= $m$ :**

In VBCI or VBPT2 calculations, the first  $m$  orbitals ( $2m$  electrons) will be frozen in the VBCI or VBPT2 calculation. In BOVB calculations, the first  $m$  orbitals will be kept as VBSCF orbitals. The default value is 0, which means all orbitals will be counted in VBCI, VBPT2 or BOVB.

**GUESS=option:**

This keyword describes the way to generate or read the initial guess for a VB computation. Valid options can be:

- AUTO: The program automatically provides guess orbitals by diagonalizing a fragment-localized Fock matrix. This is the default option.
- UNIT: The first basis function of an orbital in \$ORB is set to be the guess for the orbital.
- NBO: Initial guess will be obtained from NBOs.
- READ: Guess orbitals are read from external file, which should be provided by user.
- MO: Initial guess of VB orbitals will be obtained directly from molecular orbitals.
- RDCI: Initial guess in VBCI type will be given by users.

**NOTE:**

- a. GUESS=MO cannot be used with BPREP.
- b. GUESS=NBO cannot be used with BPREP and needs an extra preparation by NBOPREP (see 6.4).
- c. GUESS=AUTO cannot be used when some orbitals contain only one basis function (see section 4.6).

**WFNTYP=option:**

Options for the way to expand the many-electron wave functions of system.

- STR: VB structures are used. (Default)
- DET: VB determinants are used for state functions, instead of VB structures.

**VBFTYP=option:**

Options for the way to expand VB structures.

- PPD: paired-permanent-determinant algorithm is used.
- DET: Slater determinant algorithm is used.

By default, the program will decide which one to use according to the system, method, or algorithm the users choose.

**NOTE:**

1. PPD expansion can be used only with ISCF=1 or ISCF=3.
2. ISCF=5, VBPT2, VBCI, DFVB, solvation VB methods, DEN, and IPRINT $\geq$  2 will use DET expansion automatically.
3. All systems with multiplicity larger than 2 will be calculated with DET expansion.
4. Systems with electrons in VB calculation larger than 14 will be calculated with DET expansion.

**ORBTYP=option:**

Specify the type of VB orbitals.

- HAO: Hybrid Atomic Orbitals are used.
- BDO: Bond Distorted Orbitals are used.
- OEO: Overlap Enhanced Orbitals are used.
- GEN: VB orbitals are defined by users. (Default)

**NOTE:**

- \$FRAG (see Section 4.5) is needed if ORBTYP=HAO is specified. The \$FRAG section will specify the fragments based on atoms or basis functions and orbitals will be assigned in \$ORB section based on the fragment definitions in \$FRAG.
- ORBTYP=OEO does not need \$FRAG and \$ORB sections since the OEOs are delocalized in the whole system.
- ORBTYP=GEN does not need \$FRAG section, and orbitals will be described in terms of VB basis functions explicitly in \$ORB section.
- ORBTYP=HAO cannot be used with BPREP.
- ORBTYP=BDO can be used with other orbital types, such as ORBTYP=HAO as ORBTYP=HAO,BDO. ORBTYP=BDO is equivalent to ORBTYP=GEN,BDO.

**FRGTYP=option:**

Specify the type of fragments when ORBTYP=HAO.

ATOM: The fragments of system will be defined with atoms. This is the default.

SAO: The fragments of system will be defined with symmetrized atomic orbitals.

NOTE:

\$FRAG (see Section 4.5) is required if FRGTYP=SAO. For FRGTYP=ATOM, each atom is considered as a fragment if no FRAG section appears in the input file.

**4.2.2 Keywords for Computational Methods and Algorithms****HF/RHF/UHF/ROHF:**

A Hartree-Fock calculation will be proceeded. RHF/UHF/ROHF represent the restricted, unrestricted and restricted open-shell Hartree-Fock calculations respectively. When only "HF" is assigned, RHF will be proceeded when system is singlet and UHF for other cases.

**Density Functional Theory:**

A DFT calculation will be proceeded. Currently supported keywords and corresponding functionals are listed below:

LDA Functionals	
Slater	Slater exchange functional
VWN/VWN5	Vosko-Wilk-Nusair correlation functional
VWN1	Another Vosko-Wilk-Nusair correlation functional
SVWN/SVWN5	Slater + VWN5 XC functional
SVWN1	Slater + VWN1 XC functional
GGA Functionals	
B88	Beck88 exchange functional
LYP	Lee-Yang-Parr correlation functional
BLYP	Becke88 + Lee-Yang-Parr XC functional
Hybrid Functionals	
BHLYP	0.5 B88 + 0.5 HFX + LYP hybrid functional
B3LYP	Becke's 3 parameter hybrid functional, with VWN1 involved
B3LYP5	Becke's 3 parameter hybrid functional, with VWN5 involved

The users may use "R", "U", and "RO" ahead of the name of functional to specify re-

stricted, unrestricted or restricted open-shell calculations, the same as HF method. For example, "RB3LYP" will run the restricted B3LYP calculation. If only the name of functional is specified, restricted calculation will be run for singlet and unrestricted for others.

**VBSCF:**

A VB Self-Consistent Field computation is requested. This is the default method for the XMVB program.

**BOVB:**

Ask for a Breathing Orbital VB (BOVB) calculation.

**NOTE:**

- a. BOVB method cannot be used with VBCI.
- b. BOVB method is usually more difficult to converge than VBSCF. Thus, it is recommended to run a BOVB job with a good initial guess. It is recommended to run a VBSCF calculation first, followed by the BOVB calculation with optimized VBSCF orbitals as the initial guess.

**ABOVB:**

Ask for an Approximate Breathing Orbital VB (A-BOVB) calculation.

**BLW:**

Block Localized Wavefunction (BLW) method is requested. With this keyword specified, \$STR will not be read and the structure will be generated automatically. The users only need specify the type of VB orbitals (see FRGTYP and ORBTYP above).

**NOTE:**

The implementation of the BLW method in the program is not optimized. Users are recommended to run GAMESS-BLW for a BLW calculation.

**VBCIS:**

Ask for a VBCIS calculation.

**VBCISD:**

Ask for a VBCISD calculation.

**VBCIDS:**

Ask for a VBCIDS calculation.

**VBPT2:**

A VBPT2 computation will be performed.

**DFVB:**

Ask for a DFVB calculation.

**SCC:**

Size-Consistent Correction in DFVB computations will be switched on.

**VBEFP:**

Ask for a VBEFP calculation.

**VBPCM:**

Ask for a VBPCM calculation.

**VBEFPCM:**

Ask for a VBEFP/PCM calculation.

**TBVBSCF:**

Activate tensor-based VBSCF. Currently TBVBSCF is valid only when:

1. ISCF=5 is selected.
2. Structures are generated automatically with STR
3. Number of active electrons should be at least 4, in which 2 for both  $\alpha$  and  $\beta$  parts.

**VMAX=*n*:**

The maximum number of  $\sigma$  kept in Davidson diagonalization. The default value is 10. Only for TBVBSCF.

**READCOEF:**

Read a file "coef" (see 5.6) with coefficients of the first  $n$  structures as the initial guess of Davidson diagonalization. The file may be obtained from a previous TBVBSCF. Only for TBVBSCF.

**ISCF=*n*:**

ISCF specifies orbital optimization algorithm. The value  $n$  currently can be:

- 1: Numerical gradients with forward-difference approximation are used with the DFP-BFS algorithm. This is the default option of XMVB.
- 2: Analytical gradients in terms of basis functions with the L-BFGS algorithm. This algorithm involves only the first-order density matrix and is not suitable in cases displaying structures that are orthogonal to each other.
- 3: Numerical gradients with central-difference approximation are used with the DFP-BFS algorithm.
- 4: Optimization with Generalized Brillium Theory(GBT) is requested.
- 5: Analytical gradients in terms of VB orbitals with the L-BFGS algorithm. This is the most efficient algorithm so far. This option needs NAO and NAE.
- 6: VBSCF with full hessian matrix. NAO and NAE are needed for this option. This algorithm is potentially faster and more robust than ISCF=5, but it is still under development and thus is not recommended in the current version of the program.

**EIGMTHD=*option*:**

Specify the way to solve the secular equation and get the energy and coefficients. The avail-

able options can be:

**FULL:** Solving secular equation in the traditional way.

**ITER:** An iterative algorithm will be used to solving secular equation. This will be helpful for VBCI or VBSCF with large number of structures.

**WSTATE=EXP:**

Activate the state-average VBSCF calculation for ISCF=1-5. **WSTATE** may provide an array containing non-zero weights of the specific states. Following is the example for  $E = 0.5E_3 + 0.3E_5 + 0.2E_8$ :

```
$CTRL
NSTR=10 WSTATE(3)=0.5,0.0,0.3,0.0,0.0,0.2
$END
```

**SCF=n:**

Activate the Structure Combined Function (SCF) calculation containing  $n$  wavefunctions which share the same set of VB structures. The total energy is then expressed as the eigenvalue of these  $n$  states. Section \$SCF (see 4.9) is needed for the description of wavefunctions.

### 4.2.3 Keywords for Integrals

**INT=option:**

Read integrals from file or calculate them directly. The valid options can be:

**READ:** Read integrals from existing file "x1e.int" and "x2e.int". This is the default option.

**CALC:** Calculate integrals directly. Section \$GEO (see 4.10) is essential.

**LIBCINT** Integrals are calculated directly by an external library LIBCINT (project homepage: <https://github.com>)

**BASIS=basis\_set:**

Assigning the basis set when INT=CALC is requested. Basis sets are expressed the same way as Gaussian, i.e. 6-31G\*, aug-cc-pVTZ etc. The supported basis sets can be found in 6.2.

**NCHARGE=n:**

Charge of the system in current XMVB calculation. Default is 0, which means the neutral system. Positive numbers denote a cation system and negative numbers mean the system is anion. This keyword will also specify the number of electrons in current calculation, **NEL** is not needed anymore in such case.

**ERI=CD:**

This keyword activates the Cholesky decomposition for ERIs. Only valid with ISCF=5 and INT=CALC.

**CDTOL=float:**

The tolerance of Cholesky decomposition. Default is  $1 \times 10^{-10}$ . *Float* may be expressed like 1.d-6, 0.001 etc.



### 4.2.4 Keywords for Wave Function Analysis

**BOYS:**

Boys localization is requested for the final VB orbitals.

**NOTE:**

- It is strongly recommended to use this keyword for VBSCF. This makes VB orbitals easier to be interpreted and more physically meaningful.
- Boys localization is available only for VBSCF method.
- Boys localization can be only used in cases in which orbitals are separated into blocks, and there is no common basis function between blocks.

**DEN:**

First-order density matrix is requested. The result will be written to a file with extended name "den".

**OUTPUT=AIM:**

WFN file for AIM2000 program will be printed. This is available only in module distribution. A \$AIM with WFN filename is relevant for this keyword. Without \$AIM, the content of WFN file will be stored in ".dat" file of GAMESS-US.

### 4.2.5 Keywords for Symmetry

**SYMM=option:**

Options for the point group used in VB computation. Currently only Abelian groups are supported. More details are shown in Table 4.1

**NOTE:**

- The main axis of the molecule should be put along the  $z$ -axis when  $C_2$ ,  $C_{2v}$ ,  $C_{2h}$ ,  $D_2$  and  $D_{2h}$  are applied.
- This keyword cannot be used together with OEO type of VB orbitals (see **ORBTYP** keyword) and MO type of initial guess.

**IRRP=option:**

Options for the irreducible representation used in VB computation. Valid options are shown in Table 4.1.

Table 4.1: Supported point groups and irreducible representations, and corresponding options for keywords **SYMM** and **IRRP**

SYMM option	Point group	IRRP option	Irreducible representation
CS	$C_s$	AP, APP	$A', A''$
CI	$C_i$	AG, AU	$A_g, A_u$
C2	$C_2$	A, B	$A, B$
C2V	$C_{2v}$	A1, A2, B1, B2	$A_1, A_2, B_1, B_2$
C2H	$C_{2h}$	AG, AU, BG, BU	$A_g, A_u, B_g, B_u$
D2	$D_2$	A, B1, B2, B3	$A, B_1, B_2, B_3$
D2H	$D_{2h}$	AG, AU, B1G, B2G B3G, B1U, B2U, B3U	$A_g, A_u, B_{1g}, B_{2g}$ $B_{3g}, B_{1u}, B_{2u}, B_{3u}$

### 4.2.6 Keywords for Previous Version

The following keywords appear in the previous version and are not available since version 2.0. This part is important only for those who are used to the previous version.

**CIG:**

This keyword has been modified as GUESS=RDCI.

**DET:**

This keyword has been replaced by WFNTYP=DET.

**EXC:**

This keyword has been replaced by NSTATE=1.

**IOPT=*n*:**

This keyword has been replaced by keyword ISCF.

**IOUT=*n*:**

**NODIIS:**

**RHF:**

**VBXM=*n*:**

This keyword has been replaced by keywords WFNTYP and VBFTYP.

## 4.3 \$BFI Section (Required for BPREP)

The BFI section specifies how to transform primitive basis function to VB basis functions. The primitive basis functions are those used in GAMESS, Gaussian or PREINT and VB basis functions are used in XMVB. The Syntax of \$BFI section is:

\$BFI

```

NFROZ NBAS
List of frozen MOs
List of basis functions
$END

```

Here NFROZ is the number of frozen MOs and NBAS is the number of VB basis functions used in XMVB. Then frozen MOs and basis functions will be listed respectively. If there is no MO to be frozen, place a blank line there. The VB basis functions may be reordered according to how users list them. This new order will be used in ORB section for the orbital description.

Following are two examples for the BFI section. The first example comes from the HF molecule with 6-31G basis set, where basis functions are not hybridized:

```

$BFI
3 6
1 4 5
1 2 4 7 8 11
$END

```

In this example, three MOs 1, 4 and 5 are frozen and 6 VB basis functions are kept for the XMVB calculation. Primitive basis functions 3, 5, 6, 9 and 10 are removed from the list as the corresponding MOs are frozen. Note that the fourth VB basis function is the primitive basis function 7 according to the list, not primitive basis function 4 anymore.

The second example comes from the CH<sub>4</sub> molecule with 6-31G basis set, showing the orbital freezing and the hybridization of basis functions:

```

$BFI
3 8
1 3 4
1 1 1 1 1 1 3 3
1.0 1
1.0 2
1.0 4
1.0 7
1.0 8
1.0 11
0.5 12 0.5 14 0.5 16
0.5 13 0.5 15 0.5 17
$END

```

Here MOs 1, 3 and 4 are frozen and 8 hybrid VB basis functions are used in XMVB calculation. Line "1 1 1 1 1 1 3 3" specifies the number of primitive basis function in each VB basis function. Following lines specifies how the VB basis functions are hybridized. In this example, the 7-th VB basis function is a hybrid basis function composed of 3 primitive basis functions 12, 14 and 16.

## 4.4 \$STR Section

The STR section describes the information of VB structures or VB determinants if DET of CTRL section is specified. For VB structures, paired electrons, which may be lone pairs or covalent bonds, should be written first followed by unpaired electrons. The number of unpaired electrons depends on the spin multiplicity. For example: For a structure with three lone pairs (orbitals 1, 2, and 3), one covalent bond (orbitals 4 and 5), and one unpaired electron (orbital 6), the structure is expressed as,

```
1 1 2 2 3 3 4 5 6
```

For determinants, all alpha orbitals are listed first, followed by beta orbitals. For example: A determinant of alpha orbitals 1, 2, 3, 4, and 6 and beta orbitals 1, 2, 3, and 5 is expressed as

```
1 2 3 4 6 1 2 3 5
```

Note that it is strongly recommended to write the most important structure as the first one. This can avoid potential problems in VBCI.

If BOVB is specified in \$CTRL section, the program will try to convert the VB orbitals into breathing orbitals. It uses automatically different orbitals for different structures. For example: If the initial VB structures are:

```
1 1 2 3
```

```
1 1 2 4
```

```
1 1 3 5
```

The program will convert them to:

```
1 1 2 3
```

```
6 6 7 4
```

```
8 8 9 5
```

Note that the VB structures should be independent. VB structures are recommended to be written in the following orders:

### Inactive Active

where "Inactive" stands for the inactive orbitals which keep doubly occupied in all structures; "active" stands for the active orbitals whose occupation varies in the structures. The singly occupied orbitals in high-spin systems should always be put in the tail of the structures.

Following are the examples of typical bonding patterns and their corresponding \$CTRL and \$STR section examples, in which only active orbitals are labeled:

#### A. System of 2-electrons on 2-centers



S1



S2



S3

```
$CTRL
```

```
nstr=3 nmul=1
```

```
$END
```

```
$STR
```

```
1 2 ; S1
```

```
1 1 ; S2
```

```
2 2 ; S3
```

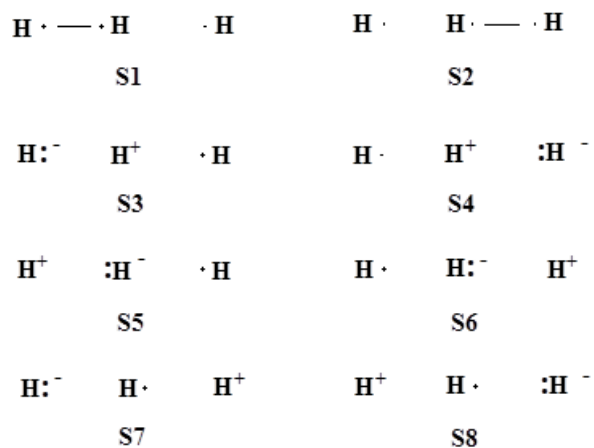
\$END

### B. System of 3-electrons on 2-centers



\$CTRL  
 nstr=2 nmul=2  
 \$END  
 \$STR  
 1 1 2 ; S1  
 2 2 1 ; S2  
 \$END

### C. System of 3-electrons on 3-centers

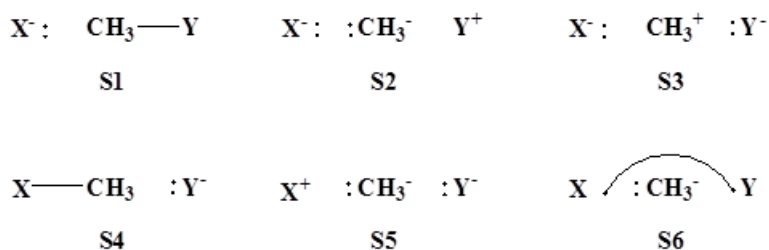


\$CTRL  
 nstr=8 nmul=2  
 \$END  
 \$STR  
 1 2 3 ; S1  
 2 3 1 ; S2  
 1 1 3 ; S3  
 3 3 1 ; S4  
 2 2 3 ; S5  
 2 2 1 ; S6  
 1 1 2 ; S7  
 3 3 2 ; S8  
 \$END

### D. System of 4-electrons and 3-centers

6 VB structures (3 VB orbitals with 4 electrons, singlet)

#### 4.5. \$FRAG SECTION (REQUIRED WHEN ORBTYP=HAO) CHAPTER 4. INPUT FILE



```

$CTRL
  nstr=6 nmul=1
$END
$STR
1 1 2 3 ; S1
1 1 2 2 ; S2
1 1 3 3 ; S3
1 2 3 3 ; S4
2 2 3 3 ; S5
2 2 1 3 ; S6
$END

```

### 4.5 \$FRAG Section (Required When ORBTYP=HAO)

Generally, the FRAG section is required if ORBTYP=HAO. In this section, fragments in which VB orbitals are localized will be defined and the orbitals will be generated with the basis functions specified in the fragments.

The syntax of \$FRAG is:

```

$FRAG
nf(1), nf(2), ... nf(N)
[basis function description(1)] lf(1,1), lf(2,1), ... lf(nf(1),1)
[basis function description(2)] lf(1,2), lf(2,2), ... lf(nf(2),2)
...
[basis function description(N)] lf(1,N), lf(2,N), ... lf(nf(N),N)
$END

```

Here the system is separated into  $N$  fragments.  $nf(i)$  means the number of atoms or basis functions in the  $i$ -th fragment, and  $lf(j, i)$  is the atom or basis function  $j$  in the  $i$ -th fragment. Basis function description is needed only when FRGTYP=SAO is chosen. Following is an example of  $\text{H}_2$  molecule with FRGTYP=ATOM:

```

$CTRL
NSTR=3 ORBTYP=HAO
FRGTYP=ATOM

```

CHAPTER 4. INPUT FILE 4.5. \$FRAG SECTION (REQUIRED WHEN ORBTYP=HAO)

```
$END
$STR
1 2
1 1
2 2
$END
$FRAG
1 1
1
2
$END
$ORB
1 1
1
2
$END
```

The above \$FRAG specifies two fragments, where one atom is in each fragment. Fragment 1 includes the first H atom and fragment 2 includes the second H atom. With this definition, users only need to specify fragment in which an orbital is located in \$ORB section. With FRGTYP=SAO, the fragments are specified by the type of basis functions. Following is an example of HF molecule with 6-31G basis set:

```
$CTRL
NSTR=3 VBFTYP=DET DEN
ISCF=5 NAO=2 NAE=2
ORBTYP=HAO FRGTYP=SAO
$END
$STR
1:4 5 6
1:4 5 5
1:4 6 6
$END
$FRAG
1 1 1 1
S 1
SPZ 2
PX 2
PY 2
$END
$ORB
1 1 1 1 1 1
2
2
3
4
```

```

2
1
$END

```

For the second fragment, "1" in the first line of \$FRAG means that the block contains basis functions located on one atom; "SPZ 2" means that the fragment includes the  $s$  and  $p_z$  basis functions in the second atom. The basis functions are described by groups of  $s$ ,  $p$ ,  $d$ ,  $f$ , etc. For example, a fragment including  $s$ ,  $p_z$ ,  $d_{xx}$ ,  $d_{yy}$ ,  $d_{zz}$ ,  $f_{zzz}$ ,  $f_{xxz}$ , and  $f_{yyz}$  basis functions in atoms 1 and 2 should be described as

```

$FRAG
2
spzdxxyzzfzzzxxzyyz 1 2
$END

```

or

```

$FRAG
2
spzdxddydzfzzzfxzfyyz 1 2
$END

```

Here "s" means basis function  $s$ , "pz" means basis function  $p_z$ , "dxyyz" means  $d_{xx}$ ,  $d_{yy}$ , and  $d_{zz}$ , and "fzzzxxzyyz" means  $f_{zzz}$ ,  $f_{xxz}$ , and  $f_{yyz}$ . The ordering of basis functions are not compulsively defined, but the basis functions with the same type of  $s$ ,  $p$ ,  $d$  and  $f$  should be written together. For example, the above description can be written equivalently as

```

$FRAG
2
spzfzzzfxzfyyzdxxdydz 1 2
$END

```

or

```

$FRAG
2
spzfxzfyyzfzzzdxxdydz 1 2
$END

```

as users like.

NOTE:

If FRGTYP=ATOM without \$FRAG, each atom is defined as one fragment. Thus, the \$FRAG section in the first example is not essential and can be erased from the file.

## 4.6 \$ORB Section (Required When ORBTYP=HAO/GEN)

The first line describes the number of basis functions (or fragments) that are used for VB orbitals. For instance,  $max(i)$  means that the  $i$ -th orbital is expanded as  $max(i)$  functions (fragments), which are specified in the following lines. If the value of  $max(i)$  is 1, it means that the  $i$ -th orbital is not optimized. From the second line, the indices of basis functions are listed,



where one orbital begins with one new line. Following is example:

```
4 4 2
3 4 5 6 ; orbital 1 is expanded with 4 basis functions (fragments)
4 3 5 6 ; orbital 2 is expanded with 4 basis functions (fragments)
1 2 ; orbital 3 is expanded with 2 basis functions (fragments)
```

NOTE:

- It is important to emphasize again that the  $n$ -th VB basis function in \$ORB section is NOT necessarily the  $n$ -th primitive basis function, but the  $n$ -th VB basis function specified in the BFI section.
- It is suggested to write the most important basis function as the first one, as the program takes the first function as the "parent" function for the orbital if GUESS=UNIT. This can avoid potential problems in convergence.
- If ORBTYP=OEO is chosen, the \$ORB is not needed. All the orbitals will be delocalized in the whole system, which means orbitals will use all basis functions.
- If the users want to freeze (not optimize) some orbitals in the calculation, simply assigning the number of basis functions (fragments) of the corresponding orbital to "0". For example, "0\*5 2 2" means that there are totally 7 VB orbitals and the first 5 will be frozen during SCF iterations. In this case, an initial guess should be provided either by "GUESS=READ" or "GUESS=MO".

## 4.7 \$AIM Section

This section is relevant if OUTPUT=AIM is specified. The content of this section is an optional file name specified by users. This file name will be used as the WFN file name. By default, the content of WFN file will be stored in ".dat" file.

## 4.8 \$GUS Section (Required when GUESS=MO/READ/RDCI)

This section describes how VB orbital obtain the initial guess. If GUESS=MO is required, the initial guess comes from MOs. An example of \$GUS from H<sub>2</sub> calculation in this situation is shown below:

```
$GUS
1 1
2 1
$END
```

The example shows that both VB orbitals 1 and 2 will get the initial guess from MO 1. All orbitals should be specified in this section.

If GUESS=READ/RDCI is required, the initial guess comes from the orbitals obtained in previous computation. In this situation, the content of \$GUS is the same as the ORB file of the previous computation. See section 5.2 for details of the content.

## 4.9 \$SCF section (Required when SCF= $n$ )

The section contains  $n$  columns of structure coefficients, each denotes a state. Following is the example for SCF=2 with 2 structures:

```
$SCF
1.0 1.0
2.0 -2.0
$END
```

Thus the result of  $\langle S_1 + 2S_2 | \hat{H} | S_1 - 2S_2 \rangle$  will be calculated, where  $S_1$  and  $S_2$  denote the 2 structures.

## 4.10 \$GEO section (Required when INT=CALC)

section contains the geometry of the system in cartesian coordinates, and the unit is Angstrom. Both Gaussian and GAMESS-US format are supported. Here both examples of the same geometry are given:

### Gaussian Format:

```
$GEO
F 0.0 0.0 -0.7
F 0.0 0.0 0.7
$END
```

### GAMESS-US Format:

```
$GEO
F 9.0 0.0 0.0 -0.7
F 9.0 0.0 0.0 0.7
$END
```

The users may choose their favorite.

# Chapter 5

## OUTPUT FILES

### 5.1 XMO FILE

The output of XMVB is stored in a file with extension "xmo". The following is an example for stand-alone XMVB:

```
*****
```

```
  M  M    MM MM    M  M    MMMM          MMMM    M
    M M    M M M    M  M    M  M          M        M
      M    M M M    M  M    MMMM  MMM  MMMM    M
    M M    M  M    M  M    M  M          M        M
  M  M    M  M    M    MMMM          MMMM  M  M
```

```
*****
```

Released on Jun 14, 2015

Cite this work as:

(a) L. Song, Z. Chen, F. Ying, J. Song, X. Chen, P. Su, Y. Mo, Q. Zhang and W. Wu, XMVB 2.1: An Ab Initio Non-orthogonal Valence Bond Program, Xiamen University, Xiamen, 361005, China, 2015 (b) L. Song, Y. Mo, Q. Zhang, W. Wu, J. Comput. Chem. 2005, 26, 514.

```
Job started at Mon Jun  8 14:36:17 2015
Work Directory at /home/fmying/VB_Workshop/Tutorial/lesson1/ex1  PID = 18199
H2 VBSCF FRAG BY ATOM
$ctrl
str=full nao=2 nae=2 # generate all VB structures with 2 active orbitals and
2 active electrons
orbtyp=hao frgtyp=atom # Construct VB orbitals with HAOs, fragmented by atom
iscf=5 # VBSCF algorithm with reduced density matrix
iprint=3 # Full print level
itmax=2000 # Maximum number of iterations is set to 2000.
$end
$frag
1 1 # 2 fragments, each contains 1 atom
```

```

1 # First fragment, containing atom 1
2 # Second fragment, containing atom 2
$end
$orb
1 1 # 2 orbitals, each contains 1 fragment
1 # Orbital 1, containing fragment 1
2 # Orbital 2, containing fragment 2
$end
-----End of Input-----
Number of 0th ion structures is:          1 from          1 to          1
Number of 0th ion determinants is:        2 from          1 to          2
Number of 1th ion structures is:          2 from          2 to          3
Number of 1th ion determinants is:        2 from          3 to          4
Total number of structures is:            3
Total number of determinants is:          4

```

READING INTEGRALS...

Reading 2-e Integrals...

Done

OPTIMIZATION METHOD: LBFGS WITH ANALYTICAL ORBITAL GRADIENT

Number of Structures: 3

The following structures are used in calculation:

```

1 ***** 1 2
2 ***** 1 1
3 ***** 2 2

```

Nuclear Repulsion Energy: 0.715104

Diagonalize Fock Matrix...

-----Initial Guess-----

```

5 5
0.5388052224 1 0.5354436033 2 0.0000000000 3 0.0000000000 4
0.1668081998 5
0.5388052224 6 0.5354436033 7 0.0000000000 8 0.0000000000 9
-0.1668081998 10

```

-----End of Guess-----

VBDET is applied

10 Coefficients 10 Independent

ITER	ENERGY	DE	GNORM
1	-1.0806051993	-1.0806051993	0.3355377748
2	-1.0995681133	-0.0189629140	0.2387997424
3	-1.1308761698	-0.0313080566	0.1539314497
4	-1.1420399230	-0.0111637531	0.1093622053
5	-1.1465763955	-0.0045364726	0.0120926244

6            -1.1466054202        -0.0000290247        0.0001989203

VBSCF converged in        6 iterations

Total Energy:            -1.14660543

First Excited:            -0.256277

The Last Change in Energy: -0.000000

Number of Iteration:     6

\*\*\*\*\* MATRIX OF OVERLAP \*\*\*\*\*

1	2	3	
1	1.000000	0.820727	0.820727
2	0.820727	1.000000	0.507832
3	0.820727	0.507832	1.000000

\*\*\*\*\* MATRIX OF HAMILTONIAN \*\*\*\*\*

1	2	3	
1	-1.857032	-1.547823	-1.547823
2	-1.547823	-1.558228	-1.080145
3	-1.547823	-1.080145	-1.558228

\*\*\*\*\* COEFFICIENTS OF STRUCTURES \*\*\*\*\*

1	0.83675	*****	1	2
2	0.09850	*****	1	1
3	0.09850	*****	2	2

\*\*\*\*\* COEFFICIENTS OF DETERMINANTS \*\*\*\*\*

a  
b

1	0.48184	*****	2
1			
2	0.48184	*****	1
2			
3	0.09850	*****	1
1			
4	0.09850	*****	2
2			

\*\*\*\*\* WEIGHTS OF STRUCTURES \*\*\*\*\*

1	0.83545	*****	1	2
2	0.08228	*****	1	1
3	0.08228	*****	2	2

Lowdin Weights

1	0.53757	*****	1	2
2	0.23121	*****	1	1
3	0.23121	*****	2	2

Inverse Weights

1	0.94072	*****	1	2
2	0.02964	*****	1	1
3	0.02964	*****	2	2

\*\*\*\*\* OPTIMIZED ORBITALS \*\*\*\*\*

	1	2
1	0.763386	0.000000
2	0.307544	0.000000
3	0.000000	0.000000
4	0.000000	0.000000
5	0.032894	0.000000
6	0.000000	0.763386
7	0.000000	0.307544
8	0.000000	0.000000
9	0.000000	0.000000
10	0.000000	-0.032894

\*\*\*\*\* ORBITALS IN PRIMITIVE BASIS FUNCTIONS \*\*\*\*\*

		1	2
1	S	0.763386	0.000000
2	S	0.307544	0.000000
3	X	0.000000	0.000000
4	Y	0.000000	0.000000
5	Z	0.032894	0.000000
6	S	0.000000	0.763386
7	S	0.000000	0.307544
8	X	0.000000	0.000000
9	Y	0.000000	0.000000
10	Z	0.000000	-0.032894

\*\*\*\*\* ORBITAL OVERLAP \*\*\*\*\*

1	2		
1	1.000000	0.712623	
2	0.712623	1.000000	

\*\*\*\*\* DENSITY MATRIX \*\*\*\*\*

1	2	3	4	5		
1	S	0.360753				
2	S	0.145336	0.058551			
3	X	0.000000	0.000000	0.000000		
4	Y	0.000000	0.000000	0.000000	0.000000	
5	Z	0.015545	0.006262	0.000000	0.000000	0.000670
6	S	0.311533	0.125507	0.000000	0.000000	0.013424
7	S	0.125507	0.050563	0.000000	0.000000	0.005408
8	X	0.000000	0.000000	0.000000	0.000000	0.000000
9	Y	0.000000	0.000000	0.000000	0.000000	0.000000
10	Z	-0.013424	-0.005408	0.000000	0.000000	-0.000578
6	7	8	9	10		
6	S	0.360753				
7	S	0.145336	0.058551			
8	X	0.000000	0.000000	0.000000		
9	Y	0.000000	0.000000	0.000000	0.000000	
10	Z	-0.015545	-0.006262	0.000000	0.000000	0.000670

ISCF = 5 currently does not support VB orbital densities

=====

XMVB ATOMIC POPULATION ANALYSIS

=====

\*\*\*\*\* POPULATION AND CHARGE \*\*\*\*\*

ATOM	MULL.POP.	CHARGE	LOW.POP.	CHARGE
1 H	1.000000	0.000000	1.000000	0.000000
2 H	1.000000	-0.000000	1.000000	-0.000000

\*\*\*\*\* ATOMIC SPIN POPULATION \*\*\*\*\*

ATOM	MULL.POP.	LOW.POP.
1 H	0.000000	0.000000
2 H	0.000000	0.000000

```

***** BOND ORDER *****
ATOM 1      ATOM 2      DIST      BOND ORDER
1 H         2 H         0.740     0.952

***** VALENCE ANALYSIS *****
ATOM      TOTAL      BONDED      FREE
VALENCE   VALENCE   VALENCE   VALENCE
1 H         1.000     0.952     0.048
2 H         1.000     0.952     0.048

***** DIPOLE MOMENT ANALYSIS *****
      DX      DY      DZ      TOTAL
0.000000  0.000000  -0.000000  0.000000

```

```

***** ENERGY DECOMPOSITION ANALYSIS *****
TOTAL VB ENERGY :      -1.146605431571
NUCLEAR REP. ENERGY :      0.715104335541
KINETIC ENERGY :      1.134053760016
POTENTIAL ENERGY :     -2.280659191588
VARIATIONAL THEOREM VALUE:      2.011067968731

```

Cpu for the Job: 0.38 (sec)

Job Finished at Mon Jun 8 14:36:17 2015

## 5.2 ORB File and GUS File

A file with extension "orb" is an output file of XMVB, which stores the optimized VB orbitals. The format is as follows:

```

max(1), max(2), ..., max(val3)
# comment for orbital 1
cvic(1,1), nvic(1,1), cvic(1,1), nvic(2,1), ..., cvic(max(1),1), nvic(max(1),1)
# comment for orbital 2
cvic(1,2), nvic(1,2), cvic(2,2), nvic(2,2), ..., cvic(max(2),2), nvic(max(2),2)
...
# comment for orbital n
cvic(1,val3), nvic(1,val3), cvic(2,val3), nvic(2, val3), ..., cvic(max(val3), val3), nvic(max(val3),
val3)

```



where  $max(i)$  stands for the number of basis functions in  $i$ -th VB orbital,  $nvic(j,i)$  is the  $j$ -th basis function in  $i$ -th VB orbital and  $cvic(j,i)$  is the coefficient of  $nvic(j,i)$ . The lines starting with "#" are treated as comments.

For VBSCF and BOVB calculations, a file of orbital guess may be provided. For VBCI calculations, the guess from a previous VBSCF calculation is required with the extension "gus". The format of orbital guess file is exactly the same as ORB file. Initial guess files with or without comments are both supported by XMVB.

NOTE:

The initial guess from previous computation with "GUESS=READ/RDCI" now is recommended to be provided in \$GUS (see section 4.8) rather than in an external file to make the computation simpler. The support for GUS file remains mainly for the compatibility.

### 5.3 XDAT File

The file with extension "xdat" is an output file of XMVB. It keeps some other information such as the orbitals in original basis form. Using utility "moldendat" can read this file and put the VB orbitals to Gaussian and GAMESS output files and Gaussian fchk files. (See Section 6.1)

### 5.4 DEN File

If keyword "den" of \$CTRL section, hybrid VB methods (DFVB, VBEFP, VBPCM and VBEFPPCM), or printing level larger than 1 are specified, one-electron density is saved to a file with "den" extension.

### 5.5 INFO File

This file stores the information of basis functions for the current system, including number of atoms, number of primitive basis functions, basis functions for each atom, and the type of each basis function. This file is essential for stand-alone distribution to carry out FRGTYP=ATOM/SAO and population analysis.

### 5.6 COEF File

This file will be obtained after a required TBVBSCF calculation. The coefficients for the structures/determinants are stored in the file and it may be used for later TBVBSCF to accelerate solving secular equation which is proceeded by Davidson Diagonalization. If the number of structures is larger than the number stored in "coef", they will be treated as coefficients of the first  $N$  structures and the rest will be set to zero.



# Chapter 6

## UTILITIES

### 6.1 Viewing VB Orbitals: Moldendat

Viewing VB orbitals is available. To do that, you need to run a utility, called "moldendat":

**moldendat.exe MOfile vmdat [denfile] >&vbfile**

where MOfile is an output file of Gaussian or GAMESS-US, or formatted Gaussian checkpoint file (.fchk); vmdat is a XMVB xdat file; if .fchk file is inputted, an optional XMVB density file with extension ".den" is also supported. The program will produce an *NEW* output file (vbfile) with the same format as input MO files, with which you can view VB orbitals with MOLDEN or MacMolPlt (for GAMESS-US only) packages.

### 6.2 Preparing Integrals: PREINT

This utility is developed to prepare integrals and MO orbitals for XMVB. To run PREINT, simply type the command as following:

**preint input >&output**

where input is the input file (see below) and output is the user-defined output file. A typical input file for F2 molecule is shown below:

```
hf cc-pVTZ libcint spher
0 1
F 0.000000 0.000000 0.000000
F 0.000000 0.000000 1.400000
```

Here keyword "libcint" means the integrals will be generated by external library LIBCINT and "spher" means spherical integrals will be generated. Both keywords are optional. Currently "spher" can only be used with "libcint".

The program provides three files: x1e.int, containing 1-e integrals and MO orbitals, x2e.int, storing 2-e integrals, and the file "INFO"(see section 5.5). You may use these files to run stand-alone XMVB jobs.

The Basis sets and elements supported by current PREINT are:

STO-2G	H-Ca,Sr
STO-3G	H-Xe
STO-6G	H-Kr
3-21G	H-Cs
3-21G*	H-Ar
3-21++G	H-Ca
3-21++G*	H-Ar
4-31G	H-Cl
6-31G	H-Zn
6-31G*	H-Kr
6-31G**	H-Zn
6-31+G	H-Ca
6-31+G*	H-Ca
6-31+G**	H-Ca
6-31++G	H-Ca
6-31++G*	H-Ca
6-31++G**	H-Ca
6-311G	H-Ca,Ga-Kr,I
6-311G*	H-Ca,Ga-Kr,I
6-311G**	H-Ca,Ga-Kr,I
6-311+G	H-Ca
6-311+G*	H-Ca
6-311+G**	H-Ca
6-311+G(2d,p)	H-Ca
6-311++G	H-Ca
6-311++G*	H-Ca
6-311++G**	H-Ca
6-311++G(2d,2p)	H-Ca
cc-pVDZ	H-Kr
cc-pVTZ	H-Ca
aug-cc-pVDZ	H-Kr
cc-pCVDZ	H-Ca
cc-pCVTZ	H-Ca
aug-cc-pCVDZ	H-Ar
aug-cc-pCVTZ	H-Ar
DZP	H-Ba,La,Hf-Rn
TZP	H-Ca

PREINT can also proceed DFT calculations. Currently supported DFT functionals are:  
Exchange functionals: Slater, B88.

Correlation functionals: VWN1, VWN5, LYP.

Exchange-correlation functionals: SVWN1, SVWN5, BLYP.

Hybrid functionals: BHHLYP, B3LYP.

More functionals will be implemented and supported in the future. All functionals support R-, U-, and RO-type calculations. To enable DFT calculations, just replace “hf” in the example with functional names, for instance

```
ub3lyp cc- pVTZ libcint spher
0 1
F 0.000000 0.000000 0.000000
F 0.000000 0.000000 1.400000
```

### 6.3 Cartesian to Spheric Integral Transformation: 6D25D

This utility transforms integrals from cartesian type to spheric (harmonic) type. Currently the utility supports D and F transformation only and not available for higher basis functions.

To run the utility, typing the command as following:

**6d25d.exe [-if gau/gms/lib] [-of gau/std]**

where option "-if" defines the sequential of cartesian F functions. Argument "gau" means the sequential in Gaussian and PREINT, "gms" means the sequential in GAMESS-US. Option and "lib" means the sequential by LIBCINT; -of" defines the output format of spheric F basis functions. Argument "gau" means the spheric F functions used in Gaussian package and "std" means standard spheric F function, which is different from the definition in Gaussian. By default, 6d25d will use Gaussian type for both input and output format.

After running 6d25d, the original cartesian integral files x1e.int, x2e.int and INFO will be overwritten by the spheric integrals. Make a backup of your cartesian integral files if you need them later.

### 6.4 Use NBOs As XMVB Initial Guess: NBOPREP

This utility read the NBOs obtained from a previous GAMESS/Gaussian calculation, and transfer them to the XMVB readable formats so that user may use them as initial guess in later XMVB calculations with keyword "GUESS=NBO".

The user need to run a GAMESS/Gaussian calculations with keyword

**\$NBO PLOT \$END**

to get files with name FILE.36 and FILE.37 which stores NBOs and PNBOs. Then run NBO-PREP as following:

**nboprep.exe outfile [NBO/PNBO]**

where "outfile" refers to the output file of GAMESS/Gaussian program, and "NBO/PNBO" tells the program which kind of NBOs should be prepared for later XMVB calculation. The user may be able to use keyword "GUESS=NBO" by copying file "orb.nbo" generated by NBO-PREP to the directory where the XMVB job will be proceeded.

### 6.5 Generate Cube file for XMVB computation: vbcubegen

This utility generates cube grid file to visualize VB orbitals with other programs. It supports module distribution or stand-alone XMVB with keyword *INT=CALC* or *INT=LIBCINT* since basis function information is essential for generating grids. The syntax of this utility is

**vbcubegen.exe xmofile**

## 6.5. GENERATE CUBE FILE FOR XMVB COMPUTATION: ~~VBCHUB~~6. UTILITIES

where *xmofile* refers to the XMO output file of the XMVB computation. After that, a cube grid file with the same file name as the xmo file will be generated, with which the user may visualize VB orbitals with programs such as GaussView, Multiwfn etc.

# Chapter 7

## TEST CALCULATIONS

### 7.1 VBSCF Calculation of HF Molecule

#### GAMESS-US INPUT FILE:

```
$contrl runtyp=energy scftyp=rhf vbtyp=xmvp $end
$basis gbasis=n21 ngauss=3 $end
$data
HF molecule 3-21G basis set
C1
H 1.0 0.000000 0.000000 0.000000
F 9.0 0.000000 0.000000 0.900000
$end
```

#### XMVB INPUT FILE:

```
HF molecule, 3 structures
$ctrl
str=full nae=2 nao=2 # automatically generate all 3 structures
iprint=3 orbtyp=hao frgtyp=sao
$end
$frag
1 1 1 1
s 1
spz 2
px 2
py 2
$end
$orb
1 1 1 1 1 1
2
2
3
4
1
2
```

\$end

NOTE:

1. VB structures are generated automatically by "STR=FULL NAO=2 NAE=2" so \$STR is not needed.
2. VB orbitals are described with SAO. See the \$FRAG and \$ORB.

## 7.2 BOVB Calculation of F<sub>2</sub> Molecule

**GAMESS-US INPUT FILE:**

```
$contrl runtyp=energy scftyp=rhf $end
$contrl vbtv=xmvp $end
$basis gbasis=n31 ngauss=6 ndfunc=1 $end
$data
F2 molecule
C1
F 9.0 0.000000 0.000000 -0.724000
F 9.0 0.000000 0.000000 0.724000
$end
```

**XMVB INPUT FILE:**

```
F2 molecule with BOVB
$ctrl
nstr=3 bovb guess=read
$end
$str
1:8 9 10
1:8 9 9
1:8 10 10
$end
$orb
8*4 3*4 8*2
1 2 5 6 9 10 11 12
16 17 20 21 24-27
1 2 5 6 9-12
16 17 20 21 24-27
3 7 14
4 8 15
18 22 29
19 23 30
1 2 5 6 9-12
16 17 20 21 24-27
$end
```

NOTE:



1. A pre-proceeded VBSHF calculation has been done and the wave function is used as the initial guess in BOVB calculation. See keyword "GUESS=READ" in \$CTRL. In this case, the initial guess must be given to XMVB, otherwise the program will exit with an error.
2. Structures and orbitals are described with abbreviations.
3. Orbitals are described with ORBTYP=GEN(default value), the ordering of the basis function is the same as in GAMESS-US.

### 7.3 VBSHF Calculation of CH<sub>3</sub>Cl Molecule

#### GAMESS-US INPUT FILE:

```
$contrl runtyp=energy scftyp=rhf $end
$contrl vbtyp=xmvp $end
$system mwords=20 $end
$basis gbasis=n31 ngauss=6 ndfunc=1 $end
$data
CH3Cl
Cl
Cl 17.0 0.000000 0.000000 0.000000
C 6.0 0.000000 0.000000 1.778500
H 1.0 -1.029975 0.000000 2.131541
H 1.0 0.514987 0.891984 2.131541
H 1.0 0.514987 -0.891984 2.131541
$end
```

#### XMVB INPUT FILE:

```
CH3Cl with 3 structures and hybrid orbitals
$ctrl
bprep nstr=3 iscf=5 nao=2 nae=2 boys
$end
$str
1:6 7 8
1:6 7 7
1:6 8 8
$end
$orb
8*4 6 8 6 8
15-22
23-30
15-22
23-30
1-6
7-14
1-6
7-14
$end
```

7.3. VBSCF CALCULATION OF CH<sub>3</sub>CL MOLECULE CHAPTER 7. TEST CALCULATIONS

---

```
$bf  
6 30  
1 2 3 4 5 6  
1 1 1 1 2 1  
1 1 1 1 2 1 3 3  
1 1 1  
1 1 1 3 3  
1 1 1  
1 1 1 2 2  
1.0 6  
1.0 9  
1.0 10  
1.0 13  
1.0 14 1.0 15  
1.0 16  
  
1.0 21  
1.0 24  
1.0 25  
1.0 28  
1.0 29 1.0 30  
1.0 31  
1.0 35 1.0 37 1.0 39  
1.0 36 1.0 38 1.0 40  
  
1.0 7  
1.0 11  
1.0 18  
  
1.0 22  
1.0 26  
1.0 33  
1.0 35 -0.5 37 -0.5 39  
1.0 36 -0.5 38 -0.5 40  
  
1.0 8  
1.0 12  
1.0 19  
  
1.0 23  
1.0 27  
1.0 34  
1.0 37 -1.0 39  
1.0 38 -1.0 40  
$end
```

NOTE:

1. BPREP is selected so that \$BFI is needed. Core orbitals are frozen and valence basis functions are hybridized.
2. A Boys localization is requested. It is recommended to the users to perform the same VB-SCF calculation without Boys localization as well and compare the difference between the orbitals of these two calculations.

## 7.4 $S_N2$ Reaction with WFN Output

**GAMESS-US INPUT FILE:**

```
$contrl runtyp=energy scftyp=rhf icharg=-1 mult=1 $end
$contrl vbtv=xmvp $end
$system mwords=20 $end
$basis gbasis=n31 ngauss=6 ndfunc=1 $end
$data
SN2 Reaction
C1
F 9.0 0.000000 0.000000 0.000000
C 6.0 0.000000 0.000000 1.783800
H 1.0 -1.076500 0.000000 1.783800
H 1.0 0.538250 0.932276 1.783800
H 1.0 0.538250 -0.932276 1.783800
F 9.0 0.000000 0.000000 3.567600
$end
```

**XMVB INPUT FILE:**

```
SN2 reaction
$ctrl
str=full nao=3 nae=4
iscf=2 output=aim
den guess=read
bprep
$end
$orb
11*6 6 8 6 6 8 6
21-31
21-31
21-31
32-42
32-42
32-42
1-6
7-14
15-20
1-6
```

7-14  
15-20  
\$end  
\$bf  
3 42  
1 2 3  
1 1 1 1 2 1  
1 1 1 1 2 1 3 3  
1 1 1 1 2 1  
1 1 1  
1 1 1 3 3  
1 1 1  
1 1 1  
1 1 1 2 2  
1 1 1  
1.0 2  
1.0 5  
1.0 6  
1.0 9  
1.0 10 1.0 11  
1.0 12  
  
1.0 17  
1.0 20  
1.0 21  
1.0 24  
1.0 25 1.0 26  
1.0 27  
1.0 31 1.0 33 1.0 35  
1.0 32 1.0 34 1.0 36  
  
1.0 38  
1.0 41  
1.0 42  
1.0 45  
1.0 46 1.0 47  
1.0 48  
  
1.0 3  
1.0 7  
1.0 14  
  
1.0 18  
1.0 22  
1.0 29  
1.0 31 -0.5 33 -0.5 35

1.0 32 -0.5 34 -0.5 36

1.0 39

1.0 43

1.0 50

1.0 4

1.0 8

1.0 15

1.0 19

1.0 23

1.0 30

1.0 33 -1.0 35

1.0 34 -1.0 36

1.0 40

1.0 44

1.0 51

\$end

\$aim

sn2.wfn

\$end

NOTE:

VBSCF calculation for the TS of SN2 reaction is performed in this example. Note that WFN file output is requested by "OUTPUT=AIM" and the file name is "sn2.wfn", as specified by \$AIM.

## 7.5 $\text{BH}_3\text{NH}_3$ Complex

**GAMESS-US INPUT FILE:**

\$contrl runtyp=energy scftyp=rhf \$end \$contrl vbtyp=xmvp \$end

\$basis gbasis=n31 ngauss=6 ndfunc=1 \$end

\$data

BH3NH3

C1

N	7.0	0.000000	0.000000	0.000000
B	5.0	0.000000	0.000000	1.664112
H	1.0	0.951736	0.000000	-0.366785
H	1.0	-0.475868	-0.824228	-0.366785
H	1.0	-0.475868	0.824228	-0.366785
H	1.0	-1.171029	0.000000	1.967388
H	1.0	0.585515	-1.014141	1.967388
H	1.0	0.585515	1.014141	1.967388

\$end

**XMVB INPUT FILE:**

```
BH3NH3
$ctrl
nstr=2 itmax=500 orbtyp=hao
$end
$stru
1:8 9 10
1:8 9 9
$end
$orb
4*10
1 3 4 5
1 3 4 5
1 3 4 5
1 3 4 5
2 6 7 8
2 6 7 8
2 6 7 8
2 6 7 8
1 3 4 5
2 6 7 8
$end
```

**NOTE:**

VBSCF calculation for BH3NH3 is proceeded in this example. Note that the orbitals are described in terms of the atoms.

## 7.6 $\text{CH}_4 + \text{SiH}_3$ H-Abstract Reaction

**GAMESS-US INPUT FILE:**

```
$contrl runtyp=energy scftyp=rohlf mult=2 $end
$contrl vbtyp=xmvp $end
$basis gbasis=n31 ngauss=6 ndfunc=1 $end
$data
CH3-H-SiH3
C1
C 6.0 0.000000 0.000000 0.000000
H 1.0 0.000000 0.000000 1.445500
Si 14.0 0.000000 0.000000 3.166000
H 1.0 1.056497 0.000000 -0.253997
H 1.0 -0.528248 -0.914953 -0.253997
H 1.0 -0.528248 0.914953 -0.253997
H 1.0 -1.399179 0.000000 3.667976
```

```
H 1.0 0.699590 1.211725 3.667976
H 1.0 0.699590 -1.211725 3.667976
$end
```

**XMVB INPUT FILE:**

```
CH3-H-SiH3 BOVB $ctrl
str=full nao=3 nae=3 nmul=2
bovb guess=read itmax=800
bprep
$end
$bf
10 28
1 2 3 4 5 6 9 10 11 12
16 17
2 5 6 9 10 11 12 37 38 39 40 41 42
23 26 27 30 31 32 33 43 44 45 46 47 48
$end
$orb
13 13 2 13 13
3-15
16-28
1 2
3-15
16-28
$end
```

**NOTE:**

TS of hydrogen abstract reaction is calculated with BOVB method. Core orbitals and  $\pi$  electrons are frozen in this case. An initial guess from a pre-proceeded VBSCF calculation should be provided.

**7.7 BDO Calculation of C<sub>2</sub>H<sub>6</sub>****GAMESS-US INPUT FILE:**

```
$contrl runtyp=energy scftyp=rhf $end
$contrl vbtyp=xmvp $end
$basis gbasis=n31 ngauss=6 ndfunc=1 $end
$data
C2H6
C1
C 6.0 0.000000 0.000000 0.000000
C 6.0 0.000000 0.000000 1.531500
H 1.0 1.023907 0.000000 -0.396508
H 1.0 -0.511953 -0.886729 -0.396508
H 1.0 -0.511953 0.886729 -0.396508
```

```

H      1.0    1.008351  0.177800  1.928008
H      1.0   -0.658155  0.784358  1.928008
H      1.0   -0.350197 -0.962158  1.928008
$end

```

**XMVB INPUT FILE:**

```

c2h6
$ctrl
nstr=1 itmax=500 orbtyp= bdo guess=read
bprep
$end
$stru
1 2 3 4 5 6 7 8 9 10 11 12 13 14
$end
$orb
14 2 14 2 14 2 14 14 14 2 14 2 14 2
1-14
15-16
1-14
17-18
1-14
19-20
1-14
21-34
21-34
35-36
21-34
37-38
21-34
39-40
$end
$bf1
2 40
1 2
2 3 4 5 6 7 8 9 10 11 12 13 14 15 31 32 33 34 35 36 17 18 19 20 21 22 23 24 25 2
6 27 28
29 30 37 38 39 40 41 42
$end

```

**NOTE:**

1. A pre-proceeded VBSCF calculation without "ORBTYPE=BDO" is essential to provide good initial guess.
2. PPD algorithm is used in this case for faster speed by keyword VBFTYPE=PPD.
3. The value of IPRINT should not be larger than 1, since PPD algorithm is not available for density calculation, which is needed for IPRINT=2 and IPRINT=3.



## 7.8 DFVB Calculation of H<sub>2</sub> + H Reaction

### GAMESS-US INPUT FILE:

```
$contrl runtyp=energy scftyp=rohf mult=2 $end
$contrl vbttyp=xmvp $end
$contrl dfttyp=lyp $end
$data
H-H-H
C1
H 1.0 0.000000 0.000000 0.000000
acct

H 1.0 0.000000 0.000000 0.931200
acct

H 1.0 0.000000 0.000000 -0.931200
acct

$end
```

### XMVB INPUT FILE:

```
H-H-H abstraction
$ctrl
str=full nmul=2 nao=3 nae=3 dfvb den
orbtyp=hao
$end
$orb
1 1 1
1
2
3
$end
```

### NOTE:

This example is a simple DFVB calculation. Structures are generated automatically and orbitals are expressed as terms of atoms.

## 7.9 VBSCF Calculation of benzene with symmetry

### GAMESS-US INPUT FILE:

```
$contrl runtyp=energy scftyp=rohf $end
$contrl vbttyp=xmvp $end
$system mwords=300 $end
$basis basnam(1)=C,C,C,C,C,C,H,H,H,H,H,H $end
$data
```

7.9. VBSCF CALCULATION OF BENZENE WITH SYMMETRY TEST CALCULATIONS

```

Benzene
C1
C 6.0 0.000000 1.395000 0.000000
C 6.0 1.208000 0.697000 0.000000
C 6.0 1.208000 -0.697000 0.000000
C 6.0 0.000000 -1.395000 0.000000
C 6.0 -1.208000 -0.697000 0.000000
C 6.0 -1.208000 0.697000 0.000000
H 1.0 0.000000 2.495000 0.000000
H 1.0 2.161000 1.247000 0.000000
H 1.0 2.161000 -1.247000 0.000000
H 1.0 0.000000 -2.495000 0.000000
H 1.0 -2.160000 -1.247000 0.000000
H 1.0 -2.160000 1.247000 0.000000

```

\$end

\$C

ccd

\$end

\$H

ccd

\$end

**XMVB INPUT FILE:**

benzene

\$ctrl

symm=d2h irrp=ag

iscf=5 frgtyp=sao orbtyp=hao

str=full nao=6 nae=6 guess=mo

\$end

\$frag

1\*12 2\*6

```

spxpydxyyzzfxxxxyxzzxyyyxyzy 1
spxpydxyyzzfxxxxyxzzxyyyxyzy 2
spxpydxyyzzfxxxxyxzzxyyyxyzy 3
spxpydxyyzzfxxxxyxzzxyyyxyzy 4
spxpydxyyzzfxxxxyxzzxyyyxyzy 5
spxpydxyyzzfxxxxyxzzxyyyxyzy 6
spxpydxyyzzfxxxxyxzzxyyyxyzy 7
spxpydxyyzzfxxxxyxzzxyyyxyzy 8
spxpydxyyzzfxxxxyxzzxyyyxyzy 9
spxpydxyyzzfxxxxyxzzxyyyxyzy 10
spxpydxyyzzfxxxxyxzzxyyyxyzy 11
spxpydxyyzzfxxxxyxzzxyyyxyzy 12
pdxzyfzzxxzyyz 1 7

```

CHAPTER 7. TEST CALCULATIONS CALCULATION OF BENZENE WITH SYMMETRY

pdxzyfzzzzxyyz 2 8  
pdxzyfzzzzxyyz 3 9  
pdxzyfzzzzxyyz 4 10  
pdxzyfzzzzxyyz 5 11  
pdxzyfzzzzxyyz 6 12

\$end

\$orb

1\*6 2\*12 1\*6

1

2

3

4

5

6

1 2

2 3

3 4

4 5

5 6

1 6

1 7

2 8

3 9

4 10

5 11

6 12

13

14

15

16

17

18

\$end

\$gus

1 1

2 1

3 1

4 1

5 1

6 1

7 7

8 8

9 7

10 8

11 7

12 8

7.9. VBSCF CALCULATION OF BENZENE WITH SYMMETRY TEST CALCULATIONS

13 9  
14 10  
15 9  
16 10  
17 9  
18 10  
19 17  
20 17  
21 17  
22 17  
23 17  
24 17  
\$end

# Appendix A

## THEORY AND METHODOLOGY

In this appendix, a brief introduction to VB theory and methodologies will be given to the users. For more detailed information, it is recommended to the users to read our reviews<sup>[8,9]</sup> and research papers.<sup>[1-7,10-21]</sup>

### A.1 Introduction to VB Theory

In quantum chemistry, the many-electron wave function for a system is expressed as a linear combination of state functions:

$$\Psi = \sum_K C_K \Phi_K \quad (\text{A.1})$$

In spin-free quantum chemistry, state functions  $\Phi_K$  should be a spin eigenfunction with anti-symmetry with respect to permutation of electron indices.<sup>[1-6]</sup> The wave function is of the form

$$\Phi_K = \hat{A}\Omega_0\Theta_K \quad (\text{A.2})$$

where  $\hat{A}$  is an antisymmetrizer,  $\Omega_0$  is an orbital product as

$$\Omega_0 = \phi_1(1)\phi_2(2)\cdots\phi_N(N) \quad (\text{A.3})$$

where  $\phi_i$  is the set of VB orbitals which can be purely localized hybrid atomic orbitals (HAOs), bond distorted orbitals (BDOs, delocalized along the bonding direction)<sup>[10]</sup>, and totally delocalized overlap enhanced orbitals (OEOs), and  $\Theta_K$  is a spin function. For VB methods, the state functions are VB functions, and their spin functions may be taken as the Rumer basis sets

$$\begin{aligned} \Theta_K &= 2^{-1/2} [\alpha(i_1)\beta(j_1) - \beta(i_1)\alpha(j_1)] \times 2^{-1/2} [\alpha(i_2)\beta(j_2) - \beta(i_2)\alpha(j_2)] \cdots \\ &= \prod_{(ij)} 2^{-1/2} [\alpha(i)\beta(j) - \beta(i)\alpha(j)] \prod_k \alpha(k) \end{aligned} \quad (\text{A.4})$$

where  $(ij)$  runs over all bonds and  $k$  over all unpaired electrons. Given an orbital product  $\Omega_0$ , a complete set of VB functions is constructed by choosing all independent spin functions  $\Theta_K$ .

The coefficients  $C_K$  in Eq. (A.1) are determined by solving the conventional secular equation  $\mathbf{HC} = E\mathbf{MC}$ , where Hamiltonian and overlap matrices are defined as follows:

$$H_{KL} = \langle \Phi_K | H | \Phi_L \rangle \quad (\text{A.5})$$

and

$$M_{KL} = \langle \Phi_K | \Phi_L \rangle \quad (\text{A.6})$$

Structural weights are given by the Coulson-Chirgwin formula

$$W_K = \sum_L C_K M_{KL} C_L \quad (\text{A.7})$$

Eqs. (A.5) and (A.6) involve  $N!$  terms due to antisymmetrizer  $\hat{A}$ . If one-electron functions are orthogonal, only a few terms are non-zero and make contributions to the matrix elements, and consequently the matrix elements can be conveniently evaluated. However, in VB methods, non-orthogonal orbitals are generally used, and thus all  $N!$  terms make contributions to the matrix elements. Although it is not necessary to expand all  $N!$  terms to evaluate a determinant, the computational demanding in VB calculations is in general much more than that in MO calculations.

## A.2 The Evaluation of Hamiltonian and Overlap Matrices

In the XMVB package, two algorithms are implemented to compute the Hamiltonian and overlap matrices: one based on the Slater determinant expansion method, and the other based on the paired-permanent-determinant method.

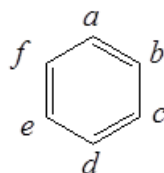
### A.2.1 Slater determinant expansion algorithm

Traditionally, an HLSP function is expressed in terms of  $2^m$  Slater determinants ( $m$  is the number of covalent bonds of structure),

$$\Phi_K = \prod_i (1 - P_i) D(\Omega_K) \quad (\text{A.8})$$

where  $D(\Omega_K)$  is a Slater determinant corresponding to Eq. (A.3),  $P_i$  is an operator that exchanges the spins of the two electrons forming the  $i$ -th bond.

**Example:** An HLSP function corresponding to a Kekulé structure of benzene is written as



$$\begin{aligned} \Phi_K = & |\bar{a}\bar{b}c\bar{d}e\bar{f}| - |\bar{a}bc\bar{d}e\bar{f}| - |\bar{a}\bar{b}c\bar{d}e\bar{f}| + |\bar{a}\bar{b}c\bar{d}e\bar{f}| \\ & - |\bar{a}\bar{b}c\bar{d}e\bar{f}| + |\bar{a}bc\bar{d}e\bar{f}| + |\bar{a}\bar{b}c\bar{d}e\bar{f}| - |\bar{a}\bar{b}c\bar{d}e\bar{f}| \end{aligned}$$

The Hamiltonian matrix element is expressed as

$$\langle D_i | H | D_j \rangle = \sum_{r,s} f_{rs} D(S_r^s) + \sum_{r \leq u, s \leq t} (g_{rs,ut} - g_{rs,tu}) D(S_{ru}^{st}) \quad (\text{A.9})$$

where  $f_{rs}$  and  $g_{rs,ut}$  are one-electron and two-electron integrals respectively, and  $D(S_r^s)$  and  $D(S_{ru}^{st})$  are the first and the second order cofactors of the overlap matrix between the two determinants respectively. Cofactors are computed by the Jacobi ratio theorem. The costs are of the order  $N^3$  for the first order and  $N^4$  for the second order cofactors at most.

## A.2.2 Paired-permanent-determinant approach

Paired-permanent-determinant (PPD) approach is based on the spin-free form of VB theory. In the spin-free VB theory, the Hamiltonian and overlap matrix elements are now written as

$$H_{KL} = \langle \Phi_K | H | \Phi_L \rangle = \sum_{P \in S_N} D_{11}^{[\lambda]}(P) \langle \Omega_K | HP | \Omega_L \rangle \quad (\text{A.10})$$

and

$$M_{KL} = \langle \Phi_K | \Phi_L \rangle = \sum_{P \in S_N} D_{11}^{[\lambda]}(P) \langle \Omega_K | P | \Omega_L \rangle \quad (\text{A.11})$$

respectively, where is the first diagonal element of the standard irreducible representation of permutation P of the symmetric group  $S_N$ . In the PPD approach, a function, called PPD, is defined as follow:

Given an  $N \times N$  square matrix  $\mathbf{A} = \{a_{ij}, i, j = 1, 2, \dots, N\}$ , the PPD of  $\mathbf{A}$  for the irreducible representation  $[\lambda]$  is the number

$$\text{ppd}(\lambda, \mathbf{A}) = \sum_{P \in S_N} D^{[\lambda]}_{11}(P) a_{1p_1} a_{2p_2} \cdots a_{Np_N} \quad (\text{A.12})$$

The evaluation of a PPD function is performed by a procedure similar to the Laplacian expansion algorithm for determinant. Hamiltonian and overlap matrix elements are computed by multiplying electronic integrals with their corresponding cofactors of PPDs. Evaluation of a PPD is more complicated than that of a determinant. But it can be beneficial when there are many bonded pairs in system. In that case there are only a few PPDs rather than numerous determinants to be evaluated.

## A.3 Orbital Optimization

The gradient vectors of energy are evaluated in four ways: the first is the numerical approximation by differential method; the second is analytical gradient based on Fock matrices, using only the first order density matrix; the third is analytical based on the first and the second order orbital density matrices; and the third is based on generalized Brillouin theorem. The first three methods are fitted for all-type orbitals, and the later one is only available for strictly localized and delocalized orbitals. The second one is suitable only when there is no orthogonality between VB functions. There are two orbital optimization methods adopted in the package. The optimization with numerical gradient is based on the Davidson-Fletcher-Powell (DFP) family of variable metric methods, and the optimization with analytical gradient is proceeded with limited-memory Broyden-Fletcher-Goldfarb-Shanno (L-BFGS)<sup>[22]</sup> method.

## A.4 The VBSCF Methods

The wave function of Valence Bond Self Consistent Field (VBSCF)<sup>[23-25]</sup> method is the linear combination of VB structures, as shown in eq. (A.1). In VBSCF method, All VB structures share the same set of VB orbitals, and both sets of the structure coefficients and VB orbitals are optimized simultaneously to minimize the total energy. This is comparable to the MCSCF method in the MO theory. VBSCF method takes care of the static electron correlation and gives equivalent results to the MO-based CASSCF calculations. It should be noted that the dynamic electron correlation is not accounted for in the VBSCF method. In XMVB, VBSCF method is the default method, thus this keyword can be ignored.

## A.5 Post-VBSCF Method

The VBSCF result includes only static correlation energy, which makes VBSCF results not accurate enough for quantitative researches. The purpose of post-VBSCF methods is to take dynamic correlation into account as much as possible to get accurate enough results. There are several post-VBSCF methods developed so far<sup>[11-17]</sup> and will be introduced in this section. It is strongly recommended to perform post-VBSCF calculations with initial guesses from a pre-proceeded VBSCF calculation. As to VBCI and VBPT2, this is enforced.

### A.5.1 The BOVB Method

The orbitals of Breathing Orbital Valence Bond (BOVB)<sup>[11-13]</sup> method are also optimized by SCF procedure, as VBSCF does. The difference between VBSCF and BOVB methods is that BOVB provides an extra degree of freedom during orbital optimization. In BOVB method, each VB structure has its own set of orbitals and are optimized independently

$$\Psi^{\text{VBSCF}} = C_1 (|\phi_a \bar{\phi}_b| - |\phi_b \bar{\phi}_a|) + C_2 |\phi_a p \bar{h} i_a| + C_3 |\phi_b \bar{\phi}_b| \quad (\text{A.13})$$

$$\Psi^{\text{BOVB}} = B_1 (|\phi_a \bar{\phi}_b| - |\phi_b \bar{\phi}_a|) + B_2 |\phi'_a p \bar{h} i'_a| + B_3 |\phi'_b \bar{\phi}'_b| \quad (\text{A.14})$$

Thus, the orbitals adopt themselves to the instantaneous field of the VB structures, rather than to the mean field of all the structures in VBSCF. This degree of freedom makes the orbitals in BOVB "Breathing" in different structures, introduces dynamic correlation, and thereby improves considerably the accuracy of VB computations.

### A.5.2 The VBCI Method

The VBCI<sup>[14,15]</sup> method is based on localized VB orbitals. In this method VB orbitals are divided to several blocks (occupied and virtual orbitals). Excited VB structures are generated by replacing occupied VB orbitals with virtual orbitals that are localized on the same block. The wave function of VBCI is the linear combination of all reference and excited VB structures

$$\Psi^{\text{VBCI}} = \sum_K \sum_i C_{Ki} \Phi_K^i \quad (\text{A.15})$$

where  $\Phi_K^i$  s CI structure coming from VBSCF structure  $K$ , including reference and excited structures, and the coefficients  $\{C_{Ki}\}$  are obtained by solving the secular equation. The VBCI



weight can be given either with eq. (A.7), which gives weights of all CI structures, or in a more compact way as

$$W_K = \sum_i W_{Ki} \quad (\text{A.16})$$

where  $W_K$  is the contracted weights of reference structure  $K$ , including weights of all CI structures coming from structure  $K$ .

Allowing for different excitations for different electronic shells, currently the VBCI method consists of the following calculation levels:

- VBCI(S,S): only single excitations are involved in either active electron or inactive electron. In brief, this is a VBCIS procedure.
- VBCI(D,S): the active shell is treated by single and double excitations, whereas the inactive shell by single excitations only. Also included in this level are double excitations which consist of a single excitation from each shell.
- VBCI(D,D): single and double excitations are involved for both active and inactive electrons, in short, VBCISD.

### A.5.3 The VBPT2 Method

Another post-VBSCF method is Valence Bond second-order Perturbation Theory (VBPT2)<sup>[16]</sup> method. The wave function of VBPT2 can be separated into 2 parts as

$$\Psi^{\text{VBPT2}} = \Psi^0 + \Psi^1 \quad (\text{A.17})$$

where VBSCF wave function is taken as the zeroth-order wave function  $\Psi^0$ , and the first-order part is the combination of singly and doubly excited wave functions

$$\Psi^1 = \sum_{R \in V^{SD}} C_R^1 \Phi_R \quad (\text{A.18})$$

To enhance the efficiency of VBPT2, the virtual orbitals are delocalized and orthogonal to the occupied space, and the excitations include all virtual orbitals. In this manner, the excited structures in VBPT2 don't belong to any fundamental structure, and the matrix elements can be calculated easily with Coulson-Slater rules.

### A.5.4 The DFVB Method

Density functional valence bond(DFVB)<sup>[17]</sup> method is a VB computational method which combines VBSCF and DFT correlation functional. In DFVB method, the wave function, density and static correlation energy is provided by VBSCF method, while the dynamic correlation energy is obtained by DFT correlation functional. The total energy formalism of DFVB is expressed as:

$$E^{\text{DFVB}} \approx E^{\text{VBSCF}} + E_C[\rho^{\text{VB}}] \quad (\text{A.19})$$

where  $E^{\text{VBSCF}}$  is the VBSCF part, including static correlation energy, and  $E_C[\rho^{\text{VB}}]$  is the DFT correlation energy obtained by a functional according to current VB density  $\rho^{\text{VB}}$ . The total energy and wave function is optimized through a SCF procedure. In practice, GGA correlation functionals, such as LYP, PW, PBE correlation functional, are recommended functionals.

## A.6 Solvation VB Methods

### A.6.1 The VBPCM Method

The VBPCM<sup>[18]</sup> method is an ab initio solvation VB method that is based on implicit solvation model PCM in which the state wave function is expressed in the usual terms as a linear combination of VB structures. The Schrödinger equation of VBPCM is expressed as

$$\left(\hat{H}^0 + \hat{V}^{\text{PCM}}\right) \Psi^{\text{VBPCM}} = E^{\text{VBPCM}} \Psi^{\text{VBPCM}} \quad (\text{A.20})$$

where  $\hat{H}^0$  is the Hamiltonian operator in vacuum and  $\hat{V}^{\text{PCM}}$  is the solvation potential obtained by PCM.<sup>[26]</sup> The VBPCM wave function and energy are optimized simultaneously in an SCF procedure. VBPCM is now available for VBSCF and BOVB.

VBPCM has been used in several researches.<sup>[19,20]</sup> VBPCM has been rewritten in XMVB 2.0 and is capable for hetero-PCM and EFP/PCM calculations.

### A.6.2 The VBEFP Method

The VBEFP<sup>[21]</sup> method is an QM/MM method in which the QM part is expressed as a VB wave function and the MM part is expressed with EFP1,<sup>[27,28]</sup> which is a polarized water model proposed by Gordon et al. The VBEFP energy is obtained by following equation

$$\left(\hat{H}^0 + \hat{V}^{\text{EFP}}\right) \Psi^{\text{VBEFP}} = E^{\text{VBEFP}} \Psi^{\text{VBEFP}} \quad (\text{A.21})$$

where  $\hat{H}^0$  is the Hamiltonian operator in vacuum and  $\hat{V}^{\text{EFP}}$  is the solvation potential obtained by EFP method. An SCF procedure is used to optimize the wave function and energy of VBEFP simultaneously.

Currently, VBEFP is available only in module distribution and is only available for VBSCF. With the use of EFP, VBEFP is very useful to take strong and short solvent-solute interactions into account.

# Appendix B

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